

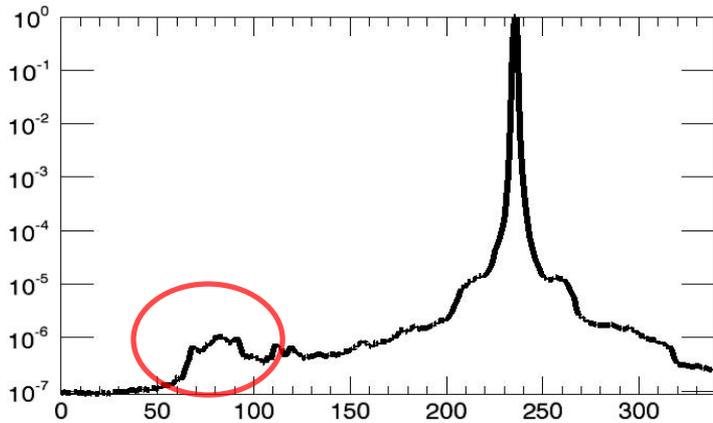
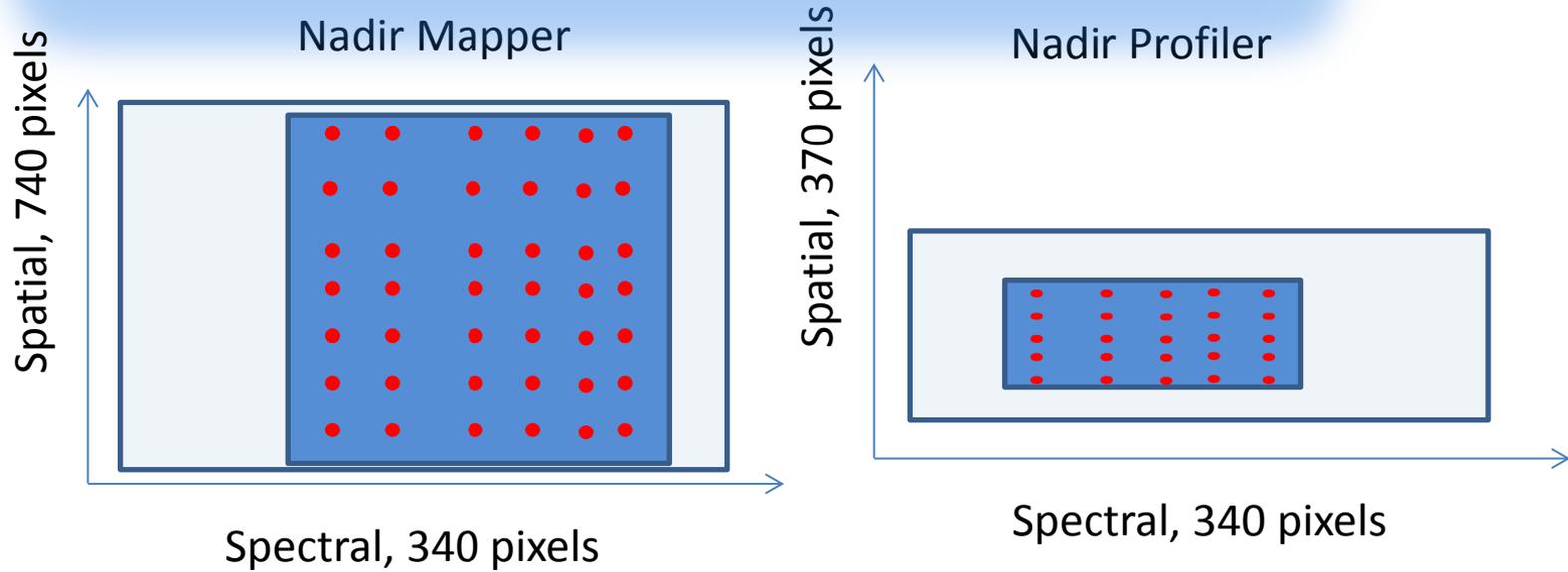
# **SNPP OMPS Nadir Instruments Stray Light Corrections**

**Hong Grace Chen, SSAI/GSFC/NASA**

**Glen Jaross, GSFC/NASA**

- **Stray Light Characterizations**
- **Stray Light Correction Approach in PEATE's APP**
- **Updates Have Been Done since Launch**
- **Validation Results and Future Plans for PEATE's Work**

# Instrument Stray Light Characterization in Pre-Launch Tests



# Stray Light Correction Approach in PEATE's APP

Simulation Method

$$C_{stray}(i_t, j_t) = \sum_{i_s} \sum_{j_s} C_{in}(i_s, j_s) PSF(i_s, j_s, i_t, j_t) + \sum_{i_s} \sum_{j_s} C_{in}(i_s, j_s) GHOST(i_s, j_s, i_t, j_t)$$

Stray Light Accuracy

1. Ground Test Measurements Accuracy

NM

||

$$C_{stray}(i_t, j_t) = \sum_{i_{s,spec}=1}^{500} \sum_{j_{s,spat}=1}^{740} C_{in}(i_{s,spec}, j_{s,spat}) Jacobian\_FullSize(i_{s,spec}, j_{s,spat}, i_t, j_t)$$

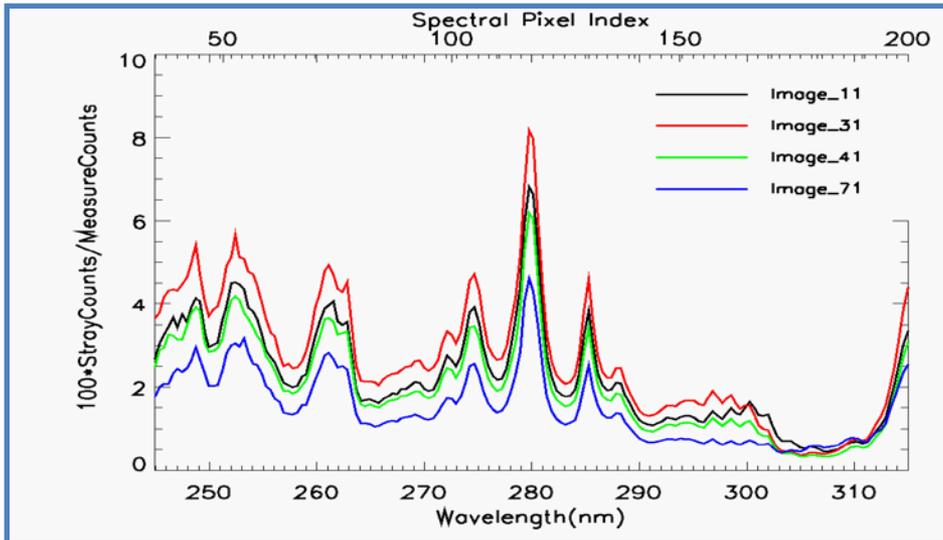
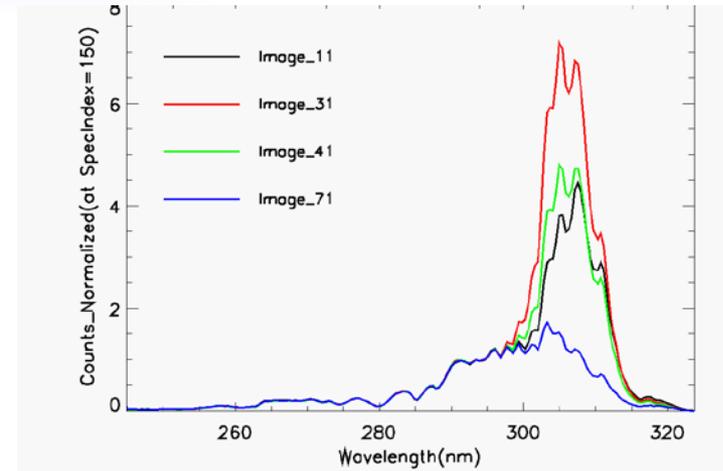
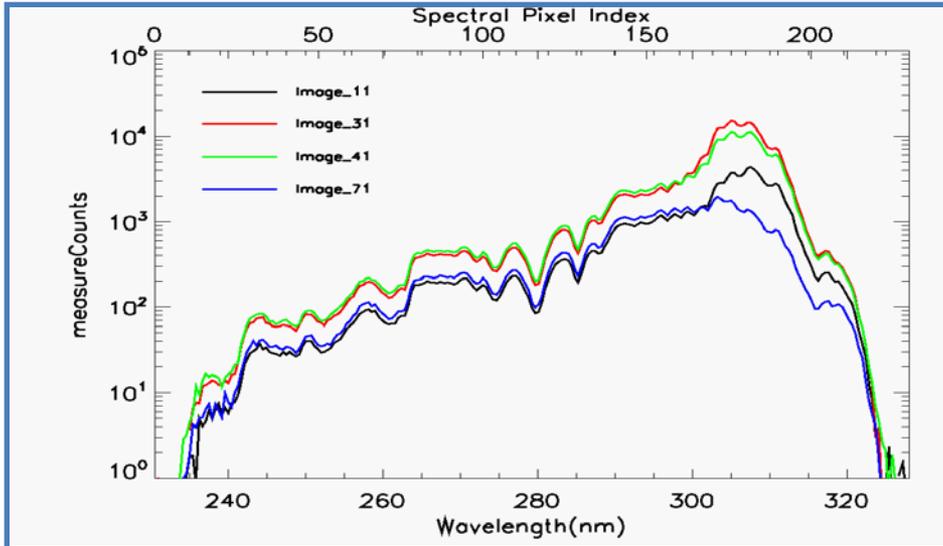
2. Source Binning Accuracy

Applied in App for NM

$$C_{stray}(i_t, j_t) = \sum_{kspec=1}^{21} \sum_{lspat=1}^{36} C_{in}(k, l) Jacobian(kspec, lspat, i_t, j_t)$$

$$C_{in}(k, l) = \sum_{m \in (k)} \sum_{n \in (l)} C_{in}(i_s, j_s)$$

# Validation Results for Nadir Profiler

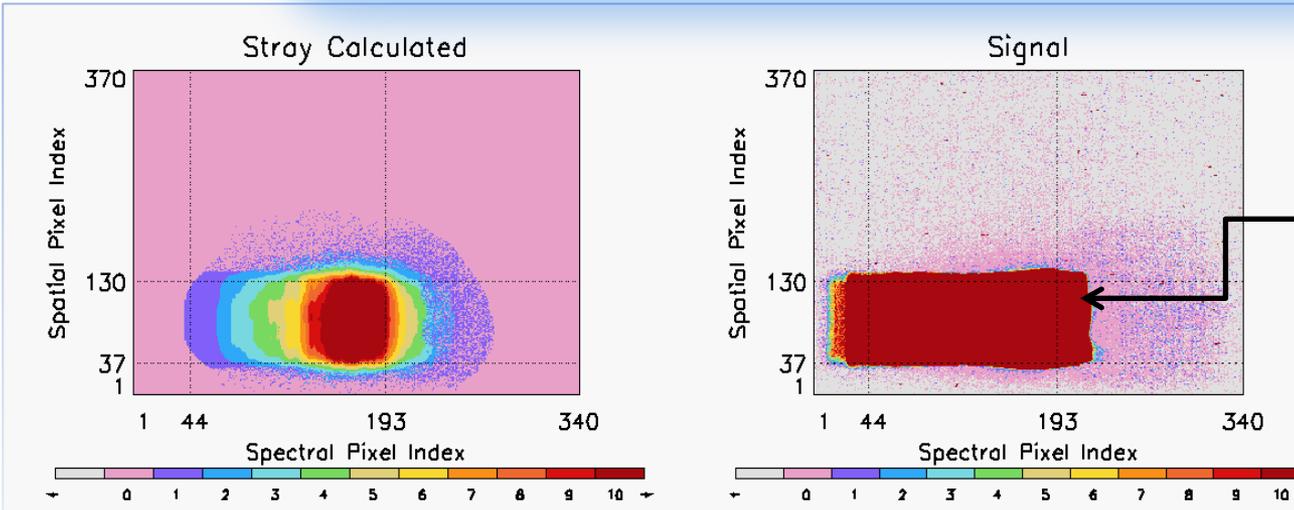


**NP full-frame Earth View  
Measurement: there are total 80  
frames in this orbit.**

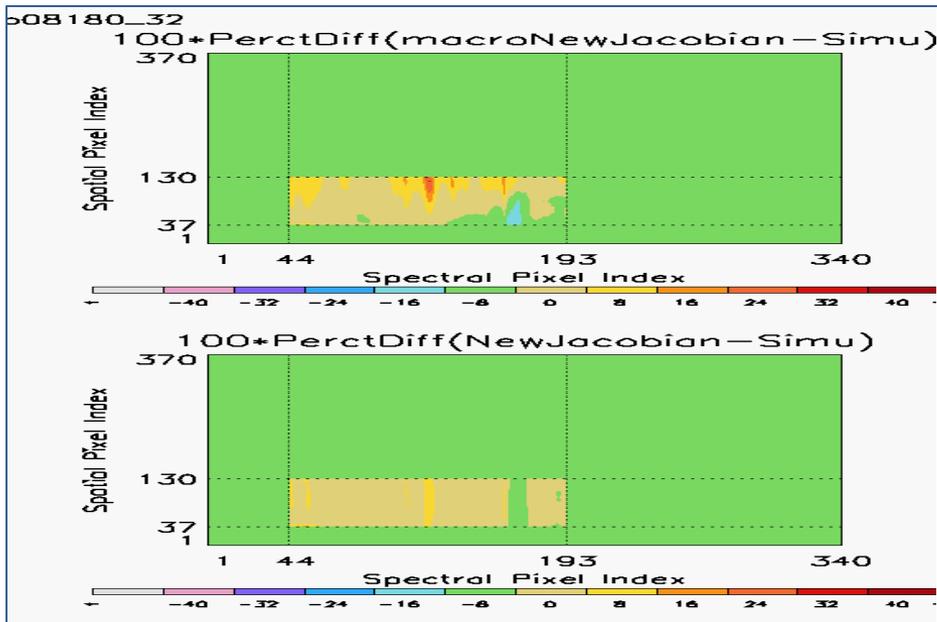
**Top Left panel: Measured counts  
(with dark and smear cleaned)**

**Lower Panel: Stray light  
percentage vs wavelength.**

# Updates since Launch and Residual for Nadir Profiler



Updates at dichroic filter transition area



IDPS Jacobian  
(SL\_COR\_COEFF  
(1,147,1,14))

PEATE Jacobian  
(370,340,14,7)

## Updates since Launch for Nadir Mapper(NM)

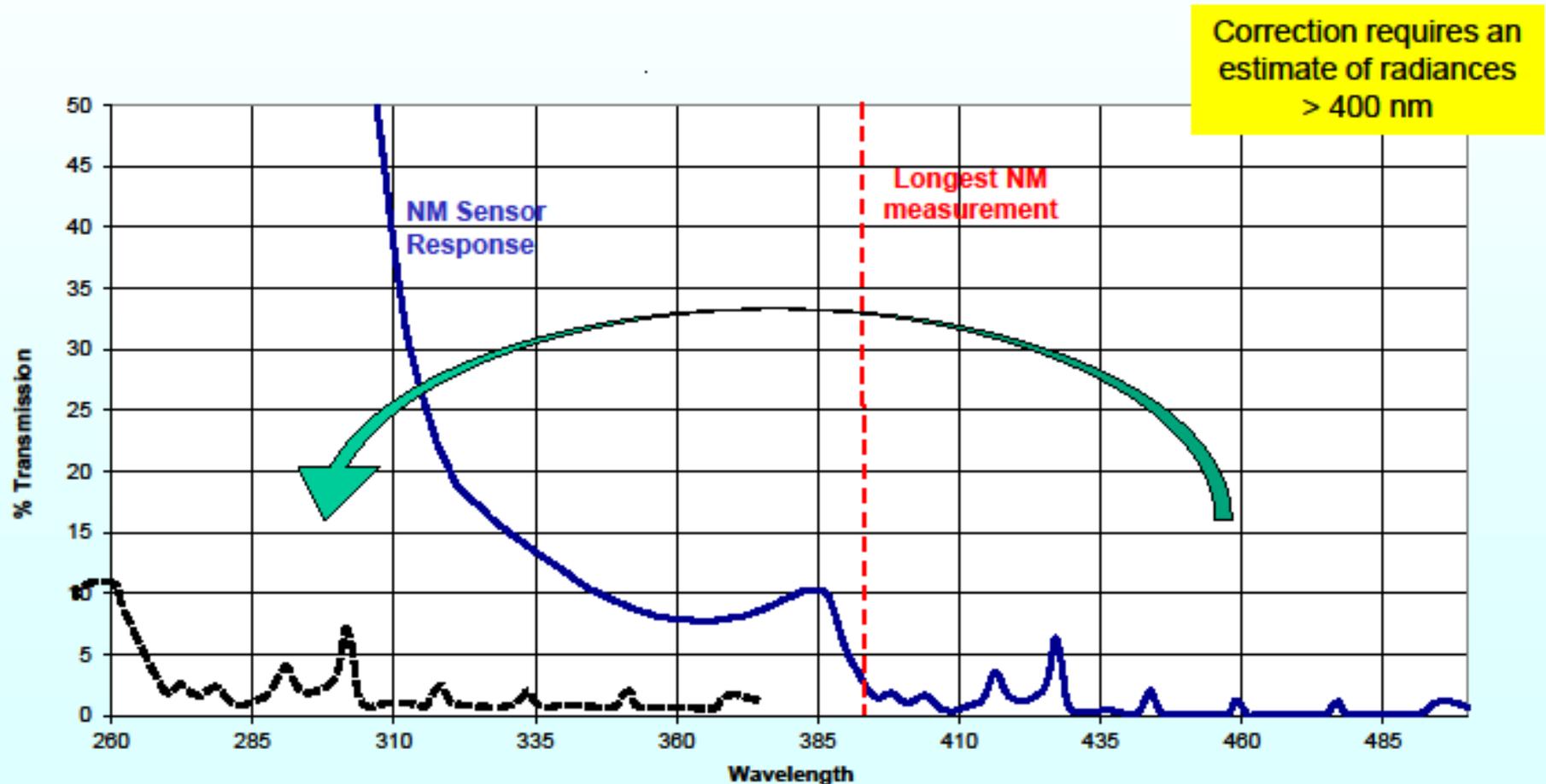
- In-band, in-field PSF Stray Light
- In-band, in-field Ghosts  
-----*First delivery to IDPS*
  
- Out-of-band PSF Stray Light
- Out-of-band Ghost (was not fully characterized during pre-launch tests)  
-----*Second delivery to IDPS*



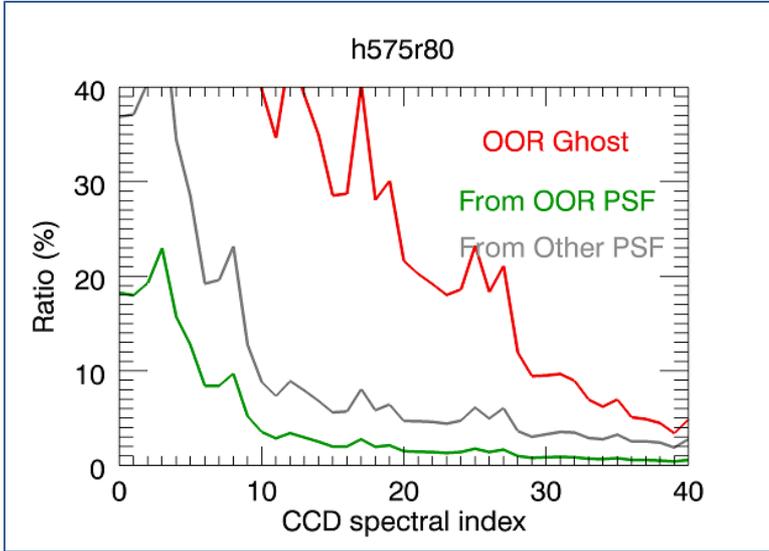
## NM ghost is a significant error source at $\lambda < 310$ nm



Reflection within NM spectrometer creates "ghost" spectrum at shorter wavelengths

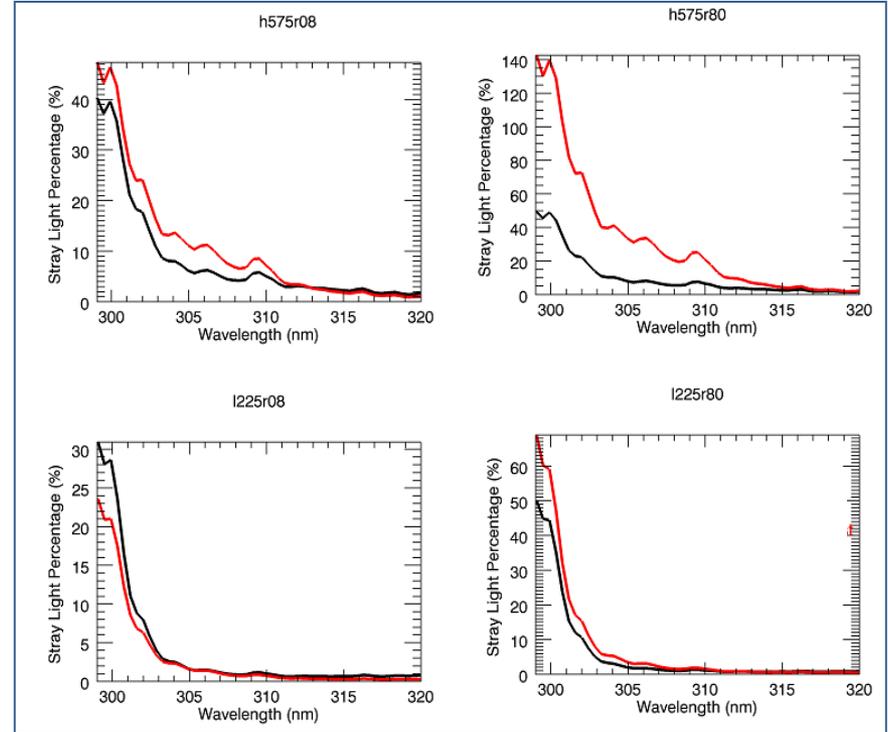


# Before Launch Simulation for NM



## Stray Light Percentage

Gray: All in-band contributions  
 Green: OOR PSF contribution  
 Red : OOR ghost contribution

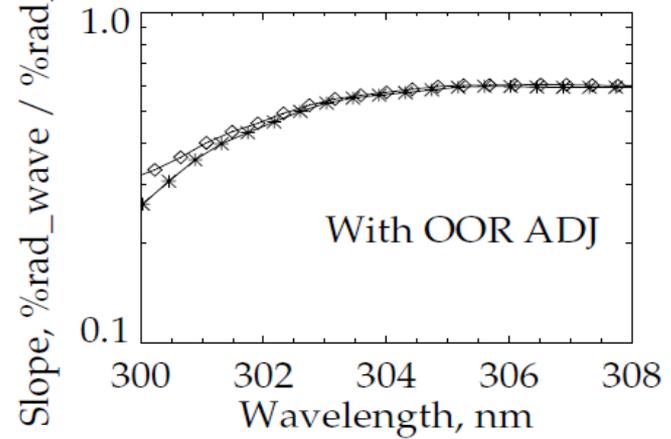
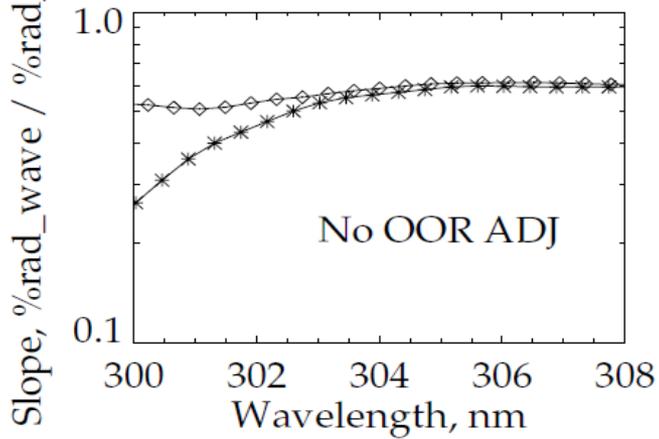
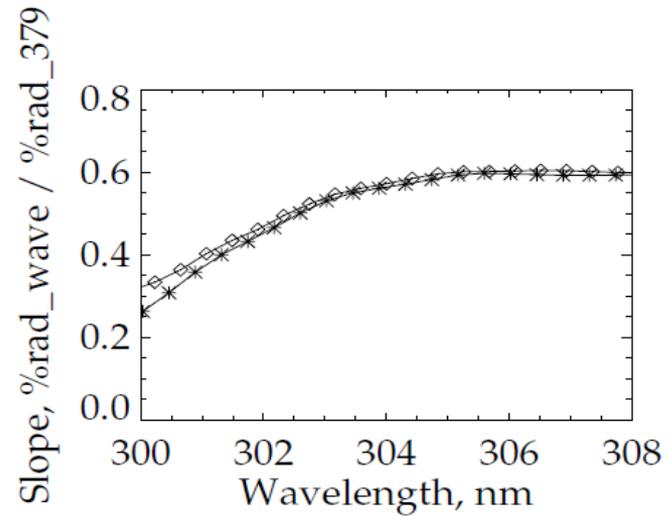
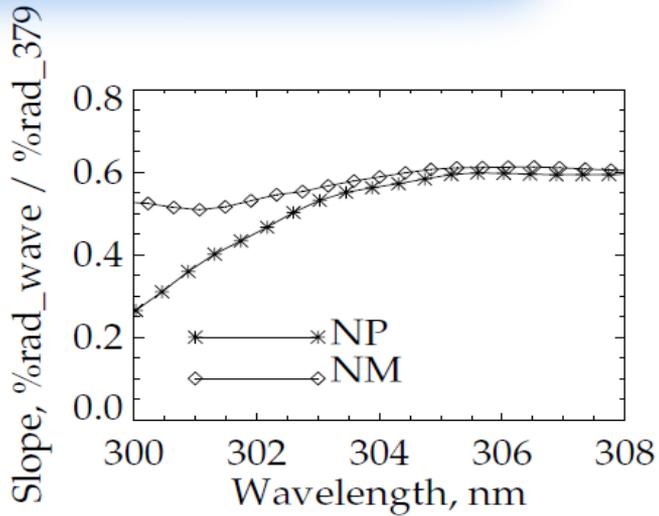


## Stray light Percentage

Black color: all in-band contributions  
 Red color : OOR contributions

# Validation Results for NM

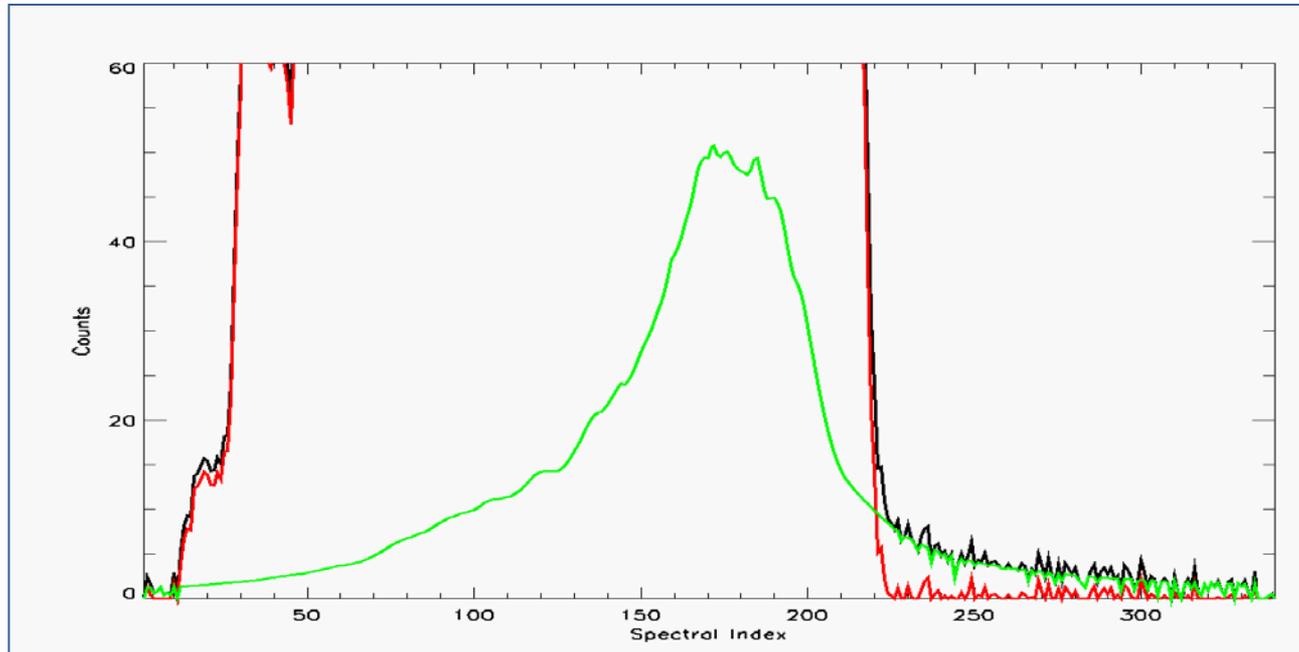
Larry F., NOAA, DR7387...



## Residual analysis: Nadir Profiler

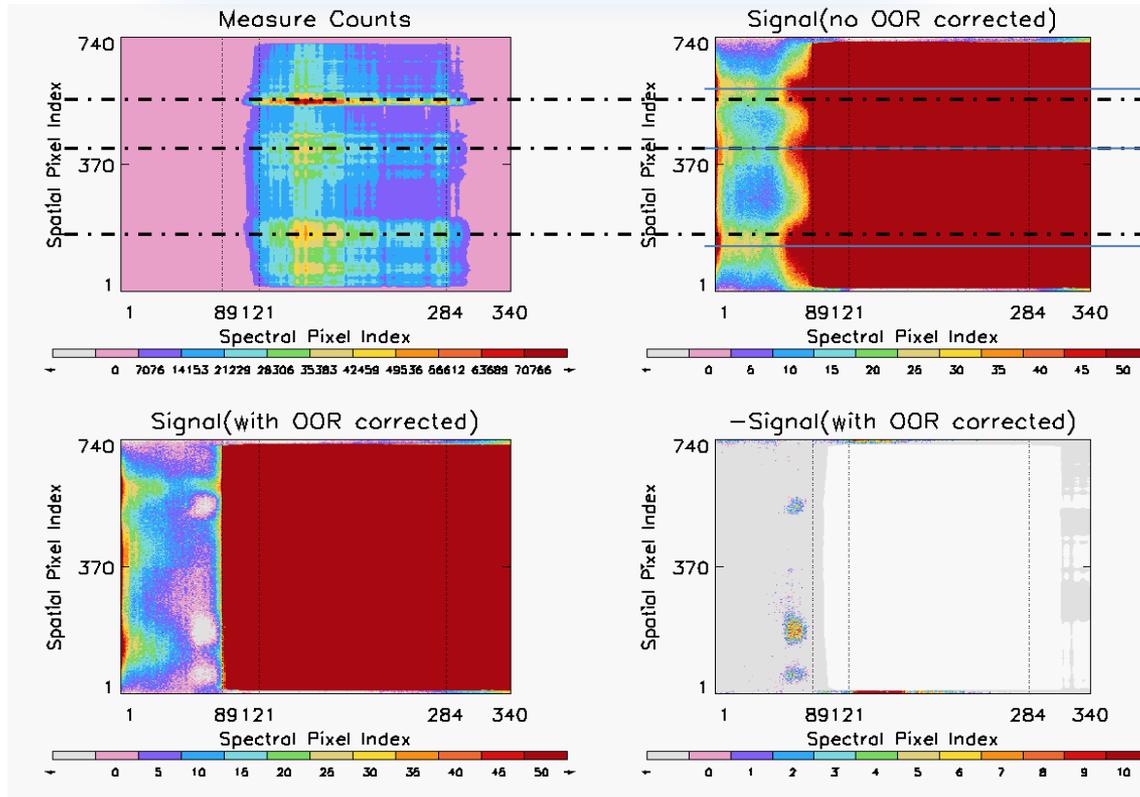
$$\text{MeasuredCounts} = \text{SignalCounts} + \text{StrayCounts}$$

? ?



**The measured counts (Black), calculated stray light counts (Green) and the signal counts (Red) vs CCD spectral index. At the longer wavelength end, the remaining signal counts are near zero; this suggests no (very few) residual stray light.**

# Residual Analysis: Nadir Mapper



NM full frame measurement before and after stray light is removed

The black-dot lines show the OOR Ghost source spatial locations; the blue thin lines show the OOR Ghost spatial locations. Clearly there are offsets....

## On-Going and Future Work for NM and NP

### **NM**

Working on OOR ghost

- \*Spatial dependence

- \*OOR source signal estimation

### **NP**

**More Need your inputs!**

Q: What the instrument imaging prosperities changes since launch?

Q: What pre-launch tests missed ?